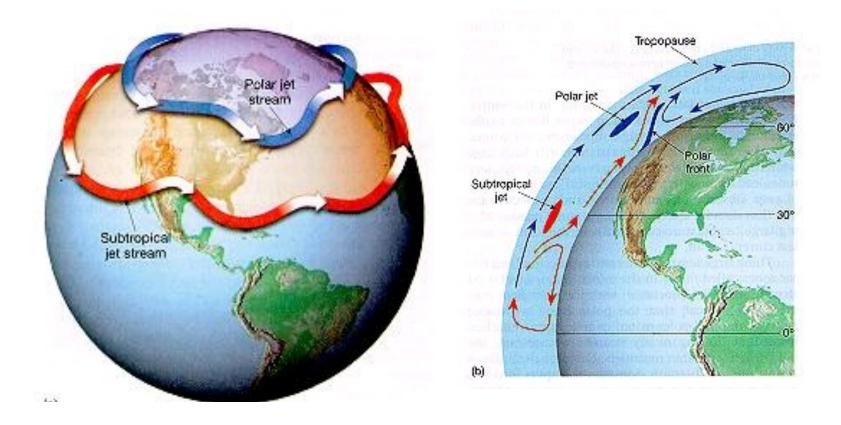
Lecture Waves

Dynamics II, 26.5.2025, 14:00

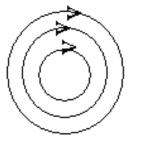
Gerrit Lohmann

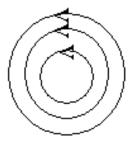
Rossby waves and the westerly wind belt



The jet stream is closely linked to the position of **Rossby waves**.

Rossby waves

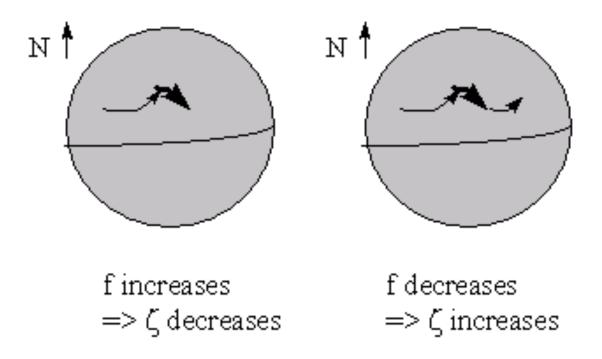




Vorticity - the tendency to spin about an axis

negative vorticity (anticyclones in NH) positive vorticity (cyclones in NH)

On the spinning Earth there is vorticity from the Earth's spin (**planetary vorticity**) and local vorticity due to cyclonic/anticyclonic behaviour (**relative vorticity**) The absolute vorticity is conserved: zeta + f = constant



Oscillations: Rossby waves

Topographic Rossby waves: standing wave fixed to a permanent forcing location

Rossby waves atmosphere

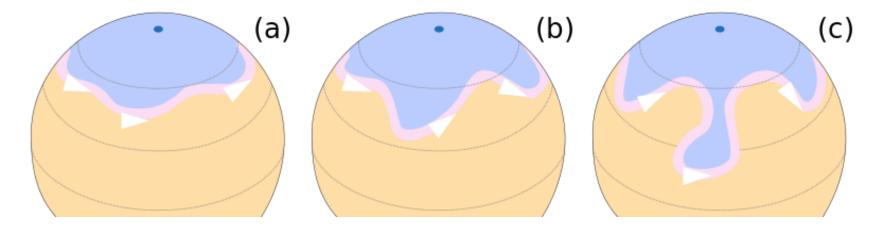


Figure 7.1: Meanders (Rossby Waves) of the Northern Hemisphere's polar jet stream developing (a), (b); then finally detaching a "drop" of cold air (c). Orange: warmer masses of air; pink: jet stream.

https://www.youtube.com/watch?v=Lg91eowtfbw&ab_channel=MetOffice-LearnAboutWeather

Waves: 2D equations

pressure in the vertically homogenous ocean is $p = g \rho (H + \eta)$

$$egin{array}{lll} rac{\partial u}{\partial t}+urac{\partial u}{\partial x}+vrac{\partial u}{\partial y}-fv&=&-grac{\partial \eta}{\partial x}\ rac{\partial v}{\partial t}+urac{\partial v}{\partial x}+vrac{\partial v}{\partial y}+fu&=&-grac{\partial \eta}{\partial y} \end{array}$$

Waves: 2D equations

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ho (H + \eta)$

$$egin{array}{lll} rac{\partial u}{\partial t}+urac{\partial u}{\partial x}+vrac{\partial u}{\partial y}-fv&=&-grac{\partial \eta}{\partial x}\ rac{\partial v}{\partial t}+urac{\partial v}{\partial x}+vrac{\partial v}{\partial y}+fu&=&-grac{\partial \eta}{\partial y} \end{array}$$

$$rac{\partial}{\partial t} \Big(
ho \left(H + \eta
ight) \Big) + rac{\partial}{\partial x} \Big(u
ho \left(H + \eta
ight) \Big) + rac{\partial}{\partial y} \Big(v
ho \left(H + \eta
ight) \Big) ~=~ 0$$

and since the density is constant it reads

$$rac{\partial}{\partial t}\eta \ + \ urac{\partial}{\partial x}\eta + vrac{\partial}{\partial y}\eta \ + rac{\partial}{\partial x}(Hu) \ + \ rac{\partial}{\partial y}(Hv) \ = \ 0$$

Waves: 2D equations

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ho \left(H + \eta
ight) \Big) + rac{\partial}{\partial x} \Big(u
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ight) \Big) + rac{\partial}{\partial y} \Big(v
ho \left(H + \eta
ight) \Big) \; = \; 0$$

and since the density is constant it reads

$$\frac{\partial}{\partial t}\eta + u\frac{\partial}{\partial x}\eta + v\frac{\partial}{\partial y}\eta + \frac{\partial}{\partial x}(Hu) + \frac{\partial}{\partial y}(Hv) = 0$$

Shallow Water Model

Linear equations $(\eta << H)$

 $egin{array}{rcl} \partial_t u&=&f\,v\,-\,g\,\partial_x\eta\ \partial_t v&=&-f\,u\,-\,g\,\partial_y\eta\ \partial_t\eta&=&-\partial_x(Hu)\,-\,\partial_y(Hv) \end{array}$

Plain waves

f and β are taken as fixed parameters in the equations. Then, the wave equations can be reduced to plain waves with eigenfunctions

 $\sim \exp(ikx + ily - i\omega t)$

In <u>fluid dynamics</u>, **dispersion** of <u>water waves</u> generally refers to <u>frequency dispersion</u>, which means that <u>waves</u> of different <u>wavelengths</u> travel at different <u>phase speeds</u>.

The <u>function</u> $\omega(k)$, which gives ω as a function of k, is known as the <u>dispersion relation</u>.

7.3.1 Inertial Waves

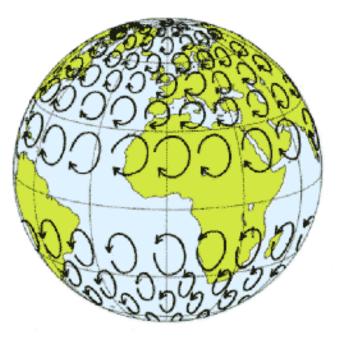
 $\partial_x \eta$, $\partial_y \eta = 0$ and $f = f_0 = const.$ (no pressure gradients and constant f)

$$egin{array}{lll} \displaystylerac{\partial u}{\partial t} - f_0 v &= 0 \ \displaystylerac{\partial v}{\partial t} + f_0 u &= 0 \end{array}$$

$$rac{\partial^2 u}{\partial t^2} = -f_0^2 u$$



Inertial waves



$$egin{array}{rcl} u(t) &=& u(0) \sin(f_0 t) \ v(t) &=& u(0) \cos(f_0 t) \end{array}$$

Figure 7.4: Schematic representation of inertial circles of air masses in the absence of other forces, calculated for a wind speed of approximately 50 to 70 m/s. Note that the rotation is exactly opposite of that normally experienced with air masses in weather systems around depressions.

Seen in drifting buoys

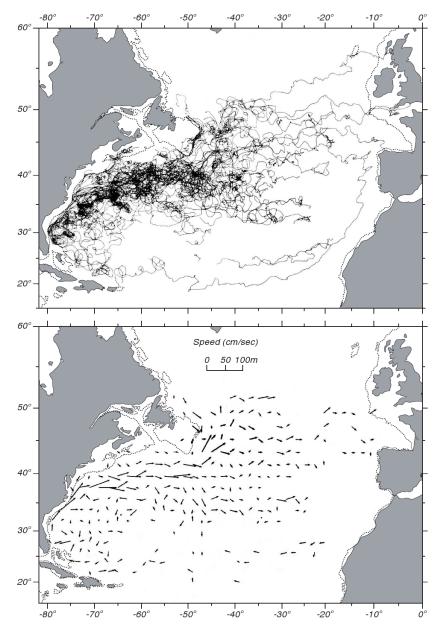
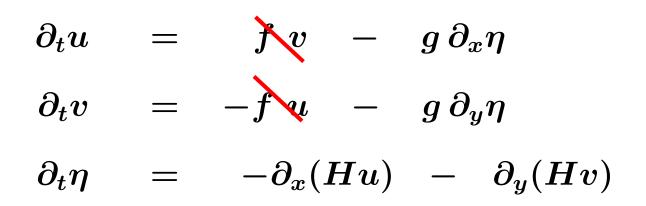


Figure 7.5: **Top:** Tracks of **110** drifting buoys deployed in the western north Atlantic. **Bottom:** Mean velocity of currents in $2^{\circ} \times 2^{\circ}$ boxes calculated from tracks above. Boxes with fewer than 40 observations were omitted. Length of arrow is proportional to speed. Maximum values are near **0.6***ms* in the Gulf Stream near **37**°N **71**°W. After Richardson (1981).

7.3.2 Gravity Waves

Shallow-water gravity waves are defined through their dynamics without the effect of the Earth's rotation, i.e. f = 0:

$$\frac{\partial^2 \eta}{\partial t^2} = g H \left(\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} \right) \eta$$
(7.28)



7.3.2 Gravity Waves

Shallow-water gravity waves are defined through their dynamics without the effect of the Earth's rotation, i.e. f = 0:

$$\frac{\partial^2 \eta}{\partial t^2} = g H \left(\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} \right) \eta$$
(7.28)

With the ansatz

$$\eta = \exp(ikx + ily - i\omega t) \tag{7.29}$$

 ω is given by

$$\omega(k,l) = \pm \sqrt{gH\left(k^2 + l^2\right)},\tag{7.30}$$

where k and l are the zonal and meridional wavenumbers. Since there is no preferred direction in the (x,y) coordinate, we simply drop the y-dependence and introduce the phase speed

$$c = \omega/k = \pm \sqrt{gH}$$
 . (7.31)

Numerical solution

Exercise 48 – Numerical solution of shallow-water gravity waves

- open shallow1D.R
- Identify the lines of the code in which the momentum equation and in which the continuum equation are solved.
- Run the program. Which type of waves do you see?
- Change the constants of water depth H, gravity g, describe your observations!
- Can you roughly estimate the phase speed of the waves?

Script of the lecture

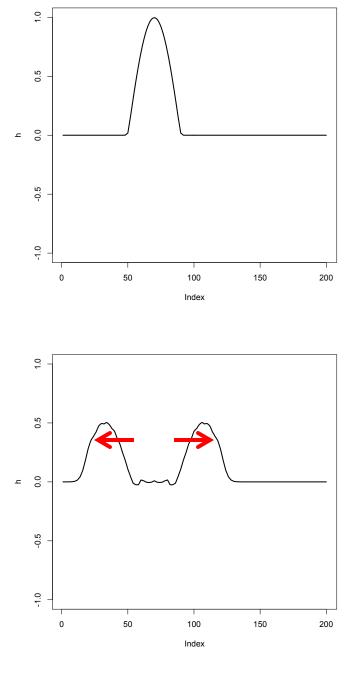


Figure 7.2: Numerical solution of 1D shallow water equation in exercise 48. Upper panel: initial condition. Lower panel: time snap shot.

Gravity wave



open shallow2D_rossby.R

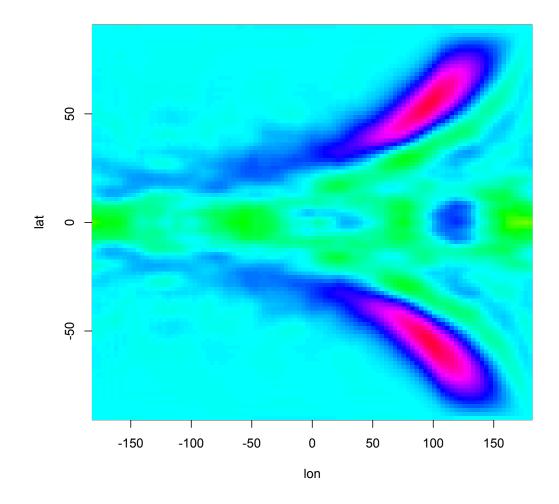
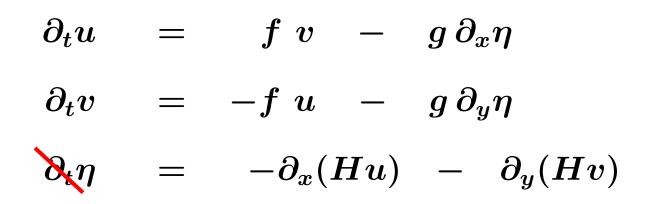


Figure 7.3: Global Rossby and Kelvin wave signatures in the exercise 49.

Shallow Water Model

Linear equations $(\eta << H)$



7.3.3 Extratropical Rossby Waves

From the equations (7.7,7.8,7.9), we drop the term $\partial_t \eta$ and introduce the stream function ψ through

$$u = \frac{\partial \psi}{\partial y} \quad ; \quad v = -\frac{\partial \psi}{\partial x}$$
 (7.36)

such that (7.9) is fulfilled. Taking $\frac{\partial}{\partial y}$ of (7.7) and subtract $\frac{\partial}{\partial x}$ of (7.8) eliminates the η term as in section 1.3:

$$\frac{\partial}{\partial t} \left(\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} \right) \psi = -\beta \frac{\partial \psi}{\partial x}$$
(7.37)

Cont. Rossby Waves

With the ansatz

$$\psi = \exp(ikx + ily - i\omega t) \tag{7.38}$$

and assumption that β is just a parameter, ω is given by

$$\omega(k,l) = -\frac{\beta k}{k^2 + l^2},\tag{7.39}$$

where k and l are the zonal and meridional wavenumbers. Again, β is used as a parameter (also called Rossby parameter) and is not expressed in terms of y:

$$\beta = \frac{df}{dy} = \frac{1}{R} \frac{d}{d\varphi} (2\Omega \sin \varphi) = \frac{2\Omega \cos \varphi}{R}$$
(7.40)

where φ is the latitude, Ω is the angular speed of the Earth's rotation, and R is the mean radius of the Earth. The wave speed $c = \omega/k = -\beta (k^2 + l^2)^{-1}$. The feature that the phase speed is faster at low latitudes can be also seen in Fig. 7.3 using the full dynamics.

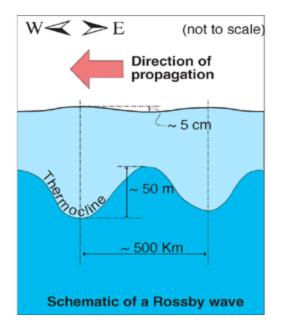
https://paleodyn.uni-bremen.de/study/Dyn2/dynamics_7.html

Rossby waves: Ocean

Existence in the oceans (<u>Carl-Gustav Rossby</u>, 1930s) has been only indirectly confirmed before the advent of satellite oceanography.

Why is it so difficult to observe them?

It is the big difference in the horizontal and vertical scale of these waves which makes them so difficult to observe. Schematic view "first-mode baroclinic" Rossby wave



speed varies with latitude and increases equatorward, order of just a few cm/s

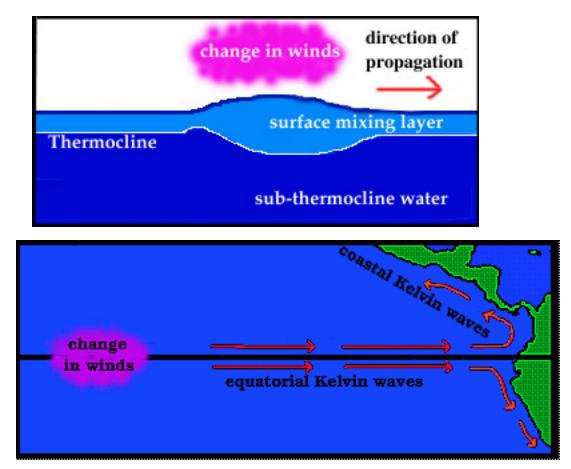
baroclinic

$$egin{array}{rcl} \partial_t u&=&f\,v\,-\,g\,\partial_x\eta\ \partial_t v&=&-f\,u\,-\,g\,\partial_y\eta\ \partial_t\eta&=&-\partial_x(Hu)\,-\,\partial_y(Hv) \end{array}$$

Exercise 51 – Baroclinic shallow-water gravity waves

In case we have a layed ocean, we consider the so-called baroclinic dynamics with the modified gravity $g' = \frac{\rho_1 - \rho_2}{\rho_1}$ using the densities $\rho_{1,2}$. Task: Derive the baroclinic dynamics using the shallow water equations for 2 different layers and subtract the equations from each other!

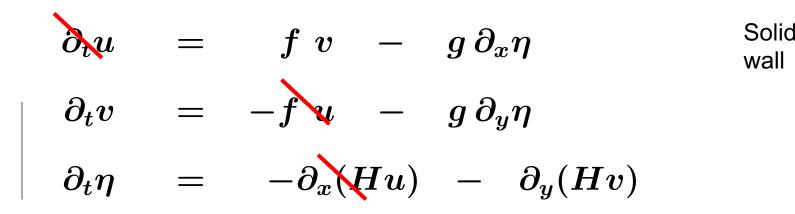
Kelvin wave



These waves, especially the surface waves are very fast moving, typically with speeds of ~ 2.8 m/s, or about 250 kilometers in a day. A Kelvin wave would take about 2 months to cross the Pacific from New Guinea to South America.

Shallow Water Model

Towards Kelvin waves



7.4 Kelvin waves

7.4.1 Coastal Kelvin waves

A Kelvin wave is a wave in the ocean or atmosphere that balances the Coriolis force against a topographic boundary such as a coastline. If one assumes that the Coriolis coefficient f is constant along the right boundary conditions, u = 0, and the zonal wind speed is set equal to zero, then the equations become the following:

$$\frac{\partial \eta}{\partial t} = -H \frac{\partial v}{\partial y}$$
(7.45)
$$\frac{\partial v}{\partial t} = -a \frac{\partial \eta}{\partial t}$$
(7.46)

$$\frac{\partial t}{\partial t} = -g \frac{\partial y}{\partial y} \tag{7.46}$$

and therefore

$$\frac{\partial^2 \eta}{\partial t^2} = gH \quad \frac{\partial^2}{\partial y^2} \eta \tag{7.47}$$

The solution to these equations yields the following phase speed: $c^2 = gH$ and $\omega = \pm cl$, which is the same speed as for shallow-water gravity waves without the effect of Earth's rotation. We see that η and v have also an x-dependence

$$\eta(x, y, t) = \tilde{\eta}(x) \exp(ily - i\omega t)$$
(7.48)

$$v(x, y, t) = \tilde{v}(x) \exp(ily - i\omega t)$$
 (7.49)

Using (7.46), we obtain

$$-i\omega \,\tilde{v}(x) = -gil \,\tilde{\eta}(x)$$
 and therefore $\tilde{v}(x) = \frac{g}{\omega} l \,\tilde{\eta}(x) = \pm \frac{g}{c} \,\tilde{\eta}(x)$ (7.50)

From the u-momentum equation
$$\frac{\partial \eta}{\partial x} = \frac{f}{g}v$$
 (7.51)

we obtain therefore
$$\frac{\partial \tilde{\eta}}{\partial x} = \pm \frac{f}{c} \tilde{\eta}$$
 (7.52)

where only the minus sign provides a useful solution (not blowing up). The solution has an expontial decay of $\tilde{\eta}(x) = \exp(-x/L_r)$ on the scale of the Rossby radius $L_r = c/f$. The wave has a trapped character along the boundary. It is important to note that for an observer traveling with

7.4.2 Equatorial Kelvin waves

Analogous we have Equatorial Kelvin waves: assume v = 0, then the equations become the following:

$$\frac{\partial \eta}{\partial t} = -H \frac{\partial u}{\partial y}$$
(7.53)
$$\frac{\partial u}{\partial t} = -g \frac{\partial \eta}{\partial y}$$
(7.54)

and therefore again

$$\frac{\partial^2 \eta}{\partial t^2} = gH \quad \frac{\partial^2}{\partial y^2} \eta \tag{7.55}$$

The solution to these equations yields the phase speed: $c^2 = gH$ and $\omega = ck$, which is the same speed as for shallow-water gravity waves without the effect of Earth's rotation. We see that η and u have also an x-dependence

7.5 Equatorial waves: Theory of Matsuno

We consider the equations (7.7, 7.8, 7.9) on the equatorial β -plane. In the equatorial region, the fluid dynamical system is described as

$$\partial_t u = \beta y v - g \partial_x \eta \tag{7.61}$$

$$\partial_t v = -\beta y \, u - g \, \partial_y \eta \tag{7.62}$$

$$\partial_t \eta = -\partial_x (Hu) - \partial_y (Hv)$$
 . (7.63)

We non-dimensionalize the system through the parameters listed in Table 7.1. In the nondimensional form (and dropping the stars in Table 7.1), the system reads then

$$\partial_t u = y v - \partial_x \eta \tag{7.64}$$

$$\partial_t v = -y \ u - \partial_y \eta \tag{7.65}$$

$$\partial_t \eta = -\partial_x u - \partial_y v$$
 . (7.66)

-> script

7.5 Equatorial waves: Theory of Matsuno

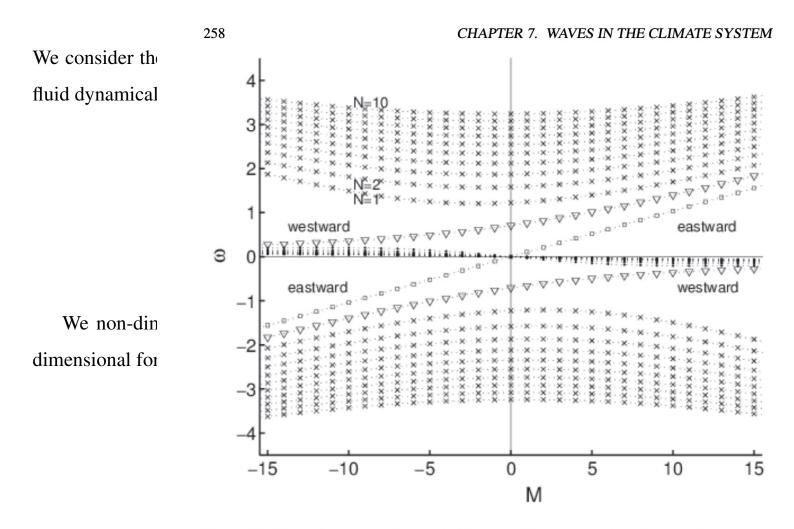


Figure 7.7: Dispersion relation for equatorial waves. Curves show dependence of frequency on zonal wave number M for mode numbers $N \leq 10$. Kelvin waves propagate eastward, Rossby waves (•) westward, while gravity waves (×) exist for both directions. Yanai waves (∇) behave Rossby-like for M < 0 and gravity-like for $M \geq 0$.

Zoo of waves

- Gravity wave
- Inertial wave
- Rossby wave
- Kelvin wave

<u>Not here:</u> Breaking waves Tsunamis Non-linear waves



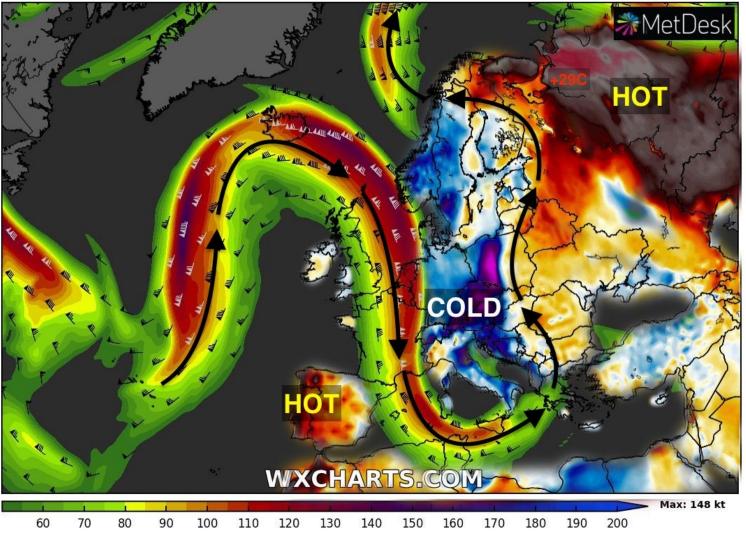


http://www.altimetry.info/thematic-use-cases/examples-of-altimetry-data-use/monitoring-el-nino-rossby-and-kelvin-waves/

Blocked waves



300 hPa Wind (kt) GFS 0.25° Run: Sat 11 May 12Z Valid: Sun 12 May 12:00 UTC

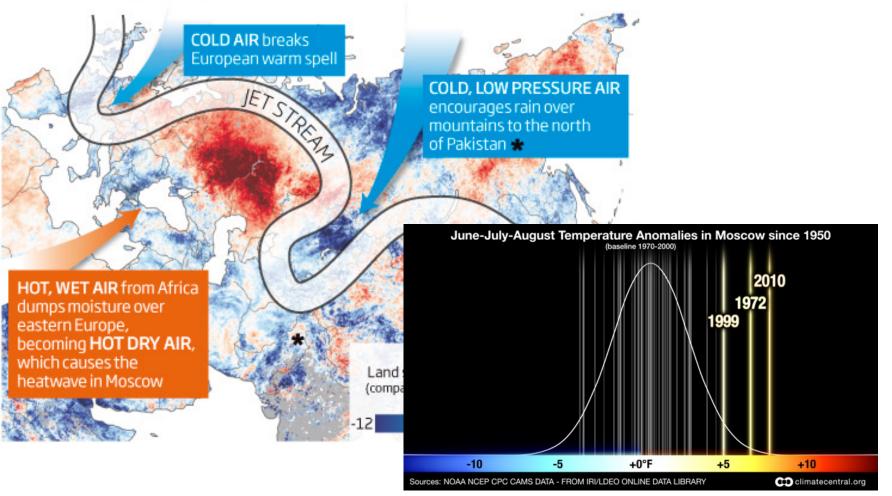


HELMHOLTZ

Blocked waves ©NewScientist

Holding pattern

In the second half of July, a blocking event froze the meanders of the jet stream over Europe and Asia. The pattern led to extreme weather across the continents



HELMHOLTZ

Atmospheric flow

Zonal Flow 13–22 Dec. 1978

Blocked Flow 10–19 Jan. 1963

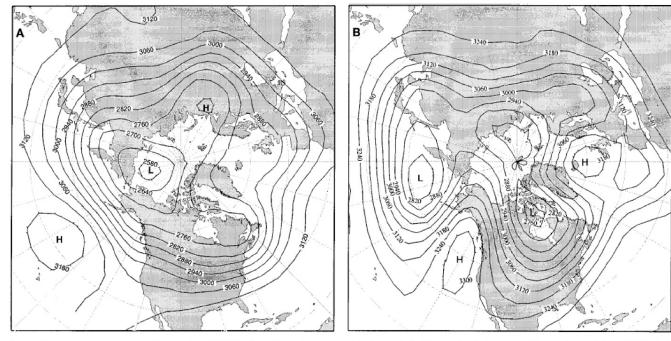


Fig. 1. Atmospheric pictures of (A) zonal and (B) blocked flow, showing contour plots of the height (m) of the 700-hPa (700 mbar) surface, with a contour interval of 60 m for both panels. The plots were obtained by averaging 10 days of twice-daily data for (A) 13 to 22 December 1978 and (B) 10 to 19 January 1963; the data are from the National Oceanic and Atmospheric

Administration's Climate Analysis Center. The nearly zonal flow of (A) includes quasi-stationary, small-amplitude waves (32). Blocked flow advects cold Arctic air southward over eastern North America or Europe, while decreasing precipitation in the continent's western part (26).

SCIENCE, Vol. 278, 28 Nov. 1997, www.sciencemag.org